

THE ROYAL ASTRONOMICAL SOCIETY OF CANADA



SASKATOON CENTRE

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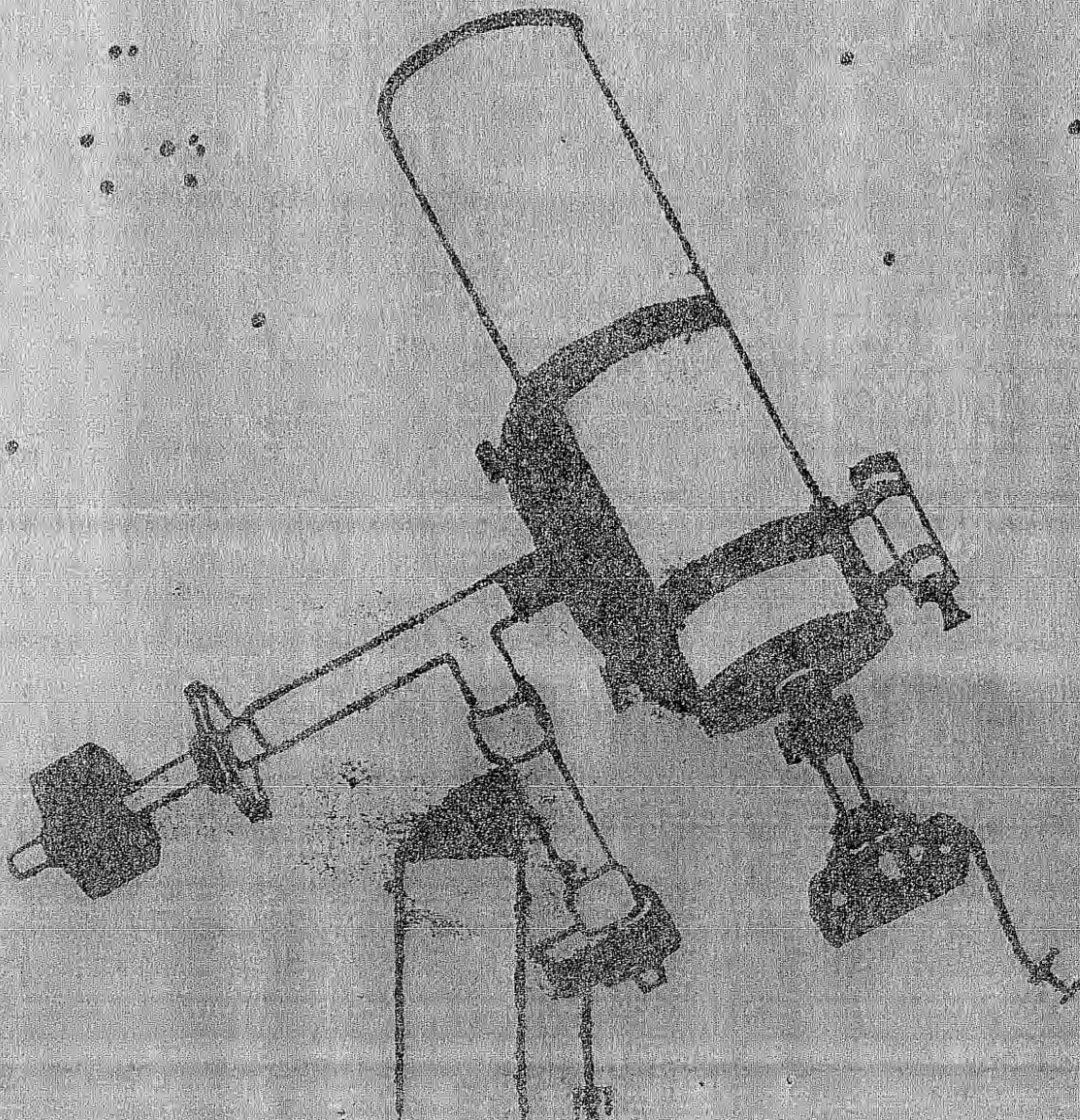
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NEWSLETTER



1(a) Telescope Aperture = 10" = 254 mm; E.F.L. = 2500 mm ; $f/\# = \frac{2500}{254} = f/9.84$
 Projection Distance, $Q_2 = 190$ mm; Ocular $f = 18$ mm

Magnification, $M = \frac{Q - f}{f} = \frac{190 - 18}{18} = 9.56X$

$f/\#$ of system = $f/\#$ of telescope $\times M = 9.84 \times 9.56 = f/94.03$

E.F.L. of system = E.F.L. of telescope $\times M = 2500 \times 9.56 = 23,900$ mm

ASA Panatomic-X = 32; with Diafine = $32 \times 6 = 192$

a)i - Venus on May 15/75 = -3.7^m = brightness, B , of 29.5 (from chart)

Calculated Exposure, $T = \frac{f^2}{ASA \times B} = \frac{94.03^2}{192 \times 29.5} = 1.56$ seconds

Using #15 filter, factor is $\times 3$; Calculated exposure = $1.56 \times 3 = 4.68$ sec.

EXPOSURE TABLE

	Calculated	With Rec. Failure
4T	18.72	140 sec
2T	9.36	59 sec
T	4.68	25.5 sec
$\frac{1}{2}T$	2.34	9.5 sec
$\frac{1}{4}T$	1.17	3.8 sec

Image Size = $\frac{\text{Obj. Size (sec)} \times \text{E.F.L.}}{200,000}$

Venus Max. = 65.56"; Image Max = 7.83 mm

Venus Min = 9.57"; Image Min = 1.14 mm

a)ii - Moon on May 17/75 - 6 days old = -4.0^m ; Brightness, B , = 32

Calculated Exposure, $T = \frac{94.03^2}{192 \times 32} = 1.44$ seconds

Using #12 filter, factor is $\times 2$; Calculated exposure = $1.44 \times 2 = 2.88$ sec.

EXPOSURE TABLE

	Calculated	With Rec. Failure
4T	11.52	77 sec
2T	5.76	32 Sec
T	2.88	13.5 sec
$\frac{1}{2}T$	1.44	5.1 sec
$\frac{1}{4}T$	0.72	1.0 sec

Moon = $30' = 1800''$

Image Size = $\frac{1800 \times 23,900}{200,000}$

= 215.0 mm

a)iii - Saturn on Mar 15/75 = 0.2^m ; $B = 12.5$; #58 filter factor = $\times 8$

Cal. Exposure = $\frac{94.03^2}{192 \times 12.5} = 3.68$ seconds $\times 8 = 29.47$ seconds.

EXPOSURE TABLE

	Calculated	With Rec. Failure
4T	117.88	1200 sec
2T	58.94	530 sec
T	29.47	240 sec
$\frac{1}{2}T$	14.74	105 sec
$\frac{1}{4}T$	7.37	44 sec

Image Size

Ball Max = 20.77"; Image = 2.48 mm
 Min = 15.07"; Image = 1.80 mm

Ring Size

Max = 48.35"; Image = 5.78 mm
 Min = 35.10"; Image = 4.19 mm

4. Color slide of total lunar eclipse with ASA 64 film, f/8 telephoto lens.
 Brilliance, B for Umbra = 0.005; for Earth Shine = 0.01

$$T_{\text{umbra}} = \frac{8 \times 8}{64 \times 0.005} = 200 \text{ seconds, plus Reciprocity Correction}$$

$$T_{\text{shine}} = \frac{8 \times 8}{64 \times 0.01} = 100 \text{ seconds, plus Reciprocity Correction}$$

Photos should be bracketted at 4T, 2T, T, $\frac{1}{2}T$ and $\frac{1}{4}T$

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5. Photo of the Pleiades. Size $2^{\circ} = 120' = 7200''$ of arc.

Usable width of 35 mm film is 24 mm.

$$\text{EFL of Telecamera System} = \frac{\text{Size} \times 200,000}{\text{Object Size}} = \frac{24 \times 200,000}{7,200} = 666.67 \text{ mm}$$

Use a 600 mm telephoto lens system.

Exposure time will depend upon usable aperture, film speed and developing, so use Tri-X, ASA 400, expose for a minimum of 10 to 20 minutes with accurate tracking, and develop film in Diafine to increase effective ASA to 2400. This should show all stars down to about 13.0 magnitude.

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6. Approximate limiting photographic magnitude for f/10 telescope with an efl of 3500 mm.

$$\text{Usable aperture} = \frac{\text{efl}}{f\#} = \frac{3500}{10} = 350 \text{ mm} = 13.78 \text{ inches}$$

$$\begin{aligned} \text{Visual Magnitude} &= 9 + 5. \log \text{Aperture in inches} = 9 + 5. \log 13.78 \\ &= 14.70 \end{aligned}$$

Photographic magnitude is about 2 to 3 magnitudes more, so the approximate photographic magnitude is 16.7^m to 17.7^m .

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NOTE: Re Question 1: Sky & Telescope gives the size of Venus as $17''$, and of Saturn - Rings $39'' \times 16.5''$, Planet $17\frac{1}{4}'' \times 15\frac{1}{2}''$.

$$\text{Image Size of Venus} = (17 \times 23,900)/200,000 = 2.03 \text{ mm}$$

$$\text{Image size of Saturn - Planet} = 2.06 \text{ mm} \times 1.85 \text{ mm}$$

$$\text{Rings} = 4.66 \text{ mm} \times 1.97 \text{ mm}$$

In Question 2, Size of Venus Image = $(17 \times 400)/200,000 = 0.034 \text{ mm}$.

1. To calculate sidereal time at Winnipeg for 7:30 pm CDST May 27/75

Change to 24 hour clock CST = 6:30 pm CST = 18:30:00 CST
 Subtract Long. Correction for WPG (29C) $\underline{-00:29:00}$ p14 Handbook
 = 18:01:00 Wpg Local Time
 Add Vernal Equinox Correction $\underline{+12:00:00}$
 = 30:01:00
 Add Apparent R.A. of Sun $\underline{+04:16:01}$ p7, Handbook
 = 34:17:01
 Subtract Sun Dial Correction $\underline{-02:58}$ p7, Handbook (Two Negatives Add)
 = 34:19:59
 Subtract whole day interval $\underline{-24:00:00}$
 Therefore, sidereal time at Winnipeg = 10:19:59 or $10^{\text{h}}19^{\text{m}}59^{\text{s}}$
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2. Star at Zenith at Winnipeg at 7:30 CST or 7:30 - 0:29 = 7:01 local Wpg time
 Time at Calgary = 6:30 MST = 6:30 - 0:36 = 5:54 local Calgary time
 Star will be at Zenith at Calgary at 7:01 - 5:54 = $1^{\text{hr}}07^{\text{min}}$ later.
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3. Sunrise & Sunset at Regina May 27/75. Regina at 50°N , Correction +58^m C
 Sunrise at 4:01am local = 4:01 + 58 = 4:58 CST
 Sunset at 19:54 local = 19:54 + 58 = 20:56 CST = 8:56 pm CST
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4. Orbital eccentricity = 0.750; perihelion distance = 0.5 a.u. = $a(1-e)$

$$\text{Mean distance, } a, = \frac{0.5}{1-e} = \frac{0.5}{1-0.75} = \frac{0.5}{0.250} = 2 \text{ a.u.}$$

$$\text{Aphelion distance} = a(1+e) = 2(1+0.75) = 2 \times 1.75 = 3.5 \text{ a.u.}$$

$$\frac{T_a^2}{T_e^2} = \frac{R_a^3}{R_e^3} ; T_a = \sqrt{\frac{2^3 \times 1^2}{1^3}} = \sqrt{8} = 2.8284 \text{ earth years} = 1,033.08 \text{ days}$$

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5. Convex lens, $f_1 = 200 \text{ mm}$; Desired $f_o = 800 \text{ mm}$ $f_o = \frac{f_1 \cdot f_2}{f_1 + f_2}$

$$\text{Therefore, } f_2 = \frac{f_o \cdot f_1}{f_1 - f_o} = \frac{800 \times 200}{200 - 800} = -266.67 \text{ mm}$$

Focal length of concave lens = -266.67 mm

5. Maximum Usable Power = 60X/inch of Aperture

$$\text{Aperture} = \frac{efl}{f/\#} = \frac{2500}{8} = 312.5 \text{ mm} = 12.3 \text{ inches}; \text{ Usable Power} = 60 \times 12.3 = 738.19X$$

$$\begin{aligned} \text{Limiting Visual Mag.} &= 9 + 5 \cdot \log(\text{Aperture in inches}) \\ &= 9 + 5 \cdot \log 12.3 = 14.45 \text{ Magnitude} \end{aligned}$$

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7. Radial Velocity of a star where λ_0 is shifted from 6640Å to 6615Å

$$\begin{aligned} \text{Velocity} = c \cdot \frac{(\lambda_b - \lambda_a)}{\lambda_0} \quad \lambda_0 &= \frac{c}{\lambda_0} = \frac{2.997925 \times 10^{10} \text{ cm/sec}}{6640 \times 10^{-8} \text{ cm}} \\ &= 4.5149472 \times 10^{14} \\ \lambda_a &= \frac{c}{\lambda_a} = \frac{2.997925 \times 10^{10} \text{ cm/sec}}{6615 \times 10^{-8} \text{ cm}} \\ &= 4.5320106 \times 10^{14} \end{aligned}$$

$$\begin{aligned} \text{Velocity} &= 2.997925 \times 10^{10} \left(\frac{4.5320106 \times 10^{14} - 4.5149472 \times 10^{14}}{4.5149472 \times 10^{14}} \right) \\ &= +1.1330097 \times 10^8 \text{ cm/sec} \\ &= +1.1330097 \times 10^3 \text{ kilometers/sec; + sign means star is nearing or} \\ &\quad \text{approaching our Solar System} \end{aligned}$$

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8. Visual magnitude = 3.44 ; Distance = 1,300 light years.

$$\text{Parallax} = \frac{10 \text{ l.y.}}{\text{Distance}} \times 0.1 = \frac{32.58 \text{ ly}}{1,300 \text{ ly}} \times 0.1 = 0.0025$$

$$\begin{aligned} \text{Absolute Magnitude, } M_v &= V + 5 + 5 \cdot \log \text{parallax} \\ &= 3.44 + 5 + 5 \cdot \log 0.0025 = -4.56497 \\ &\quad \text{or } -4.6 \end{aligned}$$

This star is Gamma Phoenix, or γ Phe, listed on p.82 in the Handbook

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9. Detail in Handbook on the star Antares. Page 89 in Handbook.

Antares - Alpha Scorpio A, R.A = 16^h 27.6^m ; Dec. = -26° 22'

Visual Magnitude = 0.92 variable; Color Index = +1.84 indicating star is red. It is a type M1, less luminous super giant with a unresolved composite spectrum Parallax is 0.019; Absolute magnitude of -5.1; distance 520 light years, with a proper motion(sideways) of 0.029 per year, and a radial velocity of -3.2 km/sec away from Earth. Star A magnitude varies from 0.86 to 1.32; Star B is 3" away from A and has a magnitude of 5.07

10. Determining Stellar Distances

- a) Trigonometric Parallax - measuring the parallax using the diameter of the Earth's orbit as a base line. Limited to relatively close stars.
- b) Spectroscopic Parallax - Intensity of spectral lines gives a relationship to the absolute magnitude. Visual magnitude can be measured and distance determined.
- c) Period Luminosity Law of Cepheid Variables - Absolute magnitude is determined from the period of variability, visual magnitude can be measured, hence distance determined.
- d) Relationship of Spectral Shift - determines radial velocity. The greater the velocity, the further away the star.

11. Motions of the Earth

- a) Daily Rotation
- b) Rotation about the Barycentre
- c) Orbital Rotation about the Sun
- d) Nutation
- e) Precession
- f) Rotation about the Galactic Center

12. NGC 6720, otherwise known as M57, the Ring Nebula in Lyra.

Located about midway between Beta Lyrae and Gamma Lyrae.

R.A = $18^{\text{h}}52^{\text{m}}5^{\text{s}}$; Dec. = $+33^{\circ}00'$; Gamma Lyrae R.A = $18^{\text{h}}49^{\text{m}}$; Dec. = $33^{\circ}20'$

Find Vega, then Gamma Lyrae. Use low powered telescope, place Gamma Lyrae in view in ocular, but below center. Stop telescope drive for $3\frac{1}{2}$ minutes and M57 should be in center of telescope.

Lyra is normally considered a Summer constellation, but may be viewed late into the Fall.

**** END ****

CORRECTION TO FUNDAMENTALS OF ASTRONOMY EXAMINATION ANSWERS

Question 1: An error was made in the calculation of the Apparent R.A. of Sun.
This should read $04^{\text{h}}12^{\text{m}}41^{\text{s}}$, changing the final result to read
 $10^{\text{h}}16^{\text{m}}39^{\text{s}}$.

Question 12: The reference star should read Beta Lyrae (R.A. $18^{\text{h}}49^{\text{m}}$, Dec $33^{\circ}20'$)
instead of Gamma Lyrae.

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RESULTS OF 1975 SASKATOON CENTRE EXAMINATIONS

Certificates of Achievement for 1975 will be awarded to the following:

FUNDAMENTALS OF ASTRONOMY

Milton AUPPERLE
Ross ELLIOTT
Dennis MARQUIS
Arthur MINERS
Debbie PALCHINSKI
Blair PETTERSON
Blair YOCHIM

ASTROPHOTOGRAPHY

Dennis MARQUIS
Debbie PALCHINSKI
Lee WARNER

ADVANCED FUNDAMENTALS OF ASTRONOMY

Gregory TOWSTEGO - Honors
Lillia WILCOX - Honors
Tory WILCOX
James YOUNG - Honors

PRACTICAL ASTROPHOTOGRAPHY

Douglas BECH - Honors
Gregory TOWSTEGO
Lillia WILCOX
Tory WILCOX
James YOUNG - Honors

Over the past three years the following members have participated in classes
and successfully passed, being awarded Certificates of Achievement:

FUNDAMENTALS: Milton AUPPERLE 75; Doug BECK 73/74; Dave DUCHALARD 73/74; Ross
ELLIOTT 75; Bruce FULTON 73/74; Ted HELSTROM 73; Alan HOLLAND 73; Hugh HUNTER 73;
Dennis MARQUIS 75; Danny McCLEAN 73; Merlyn MELBY 74; Arthur MINERS 75; Debbie
PALCHINSKI 75; Blair PETTERSON 75; Dave PRISTUPA 73/74; Greg TOWSTEGO 73/74/75;
Lee WARNER 74; Lillia WILCOX 75; Tory WILCOX 75; Jim YOUNG 74/75; Blair YOCHIM 75.
ASTROPHOTOGRAPHY: Doug BECK 73/74/75; Alan HOLLAND 73; Hugh HUNTER 73; Dennis
MARQUIS 75; Merlyn MELBY 74; Debbie PALCHINSKI 75; Dave PRISTUPA 73/74; Greg
TOWSTEGO 73/74/75; Lee WARNER 75; Lillia WILCOX 75; Tory WILCOX 75; Jim YOUNG 74/75.

All these members are to be congratulated for their personal efforts to
improve their knowledge of astronomy and/or astrophotography.

*** TELESCOPE PARTS FOR SALE ***

6 inch mirror grinding kit

6 inch mirror mount

1/8 wave diagonal for 6" reflector telescope

60 inch long aluminum tube suitable for 6" telescope

Eyeiece mount

7 X finder scope, some amateur telescope making books.

Smithsonian Astrophysical Observatory Atlas, 1969. 18" X 22"

All of the items listed above are for sale, and the seller will take \$ 100.00 for the whole lot. A very good buy! If you are interested, call Edwin Serack at 382 - 8024.

OVERDUE BOOKS

The following books are overdue. Will everyone listed please return them as soon as possible.

Popular Star Atlas by R.M.G. Inglis Blair Petterson

Astronomy Made Simple by M.H. Degani Dennis Marquis

Astronomy with Binoculars by J. Muirden Debbie Palchinski

The Telescope and the World of Astronomy by M.F. Reimer Debbie Palchinski

Amateur Astronomer by Patrick Moore Blair Petterson

Foundations of Astronomy by W.M. Smart Adrian Arbus

The Universe Nature Library Lorne Gaska

2001: A Space Odyssey by Arthur C. Clarke Richard Shaver

The Stars - Steppingstones into Space by Irving Adler . David Pristupa

Galaxies and Cosmology by P.W. Hodge Doug Beck

GENERAL MEETING

DATE: Tuesday, 17 June, 1975

TIME: 8:00 pm

PLACE: Room B111, Health Sciences Building, U of S
(across from observatory)

PROGRAM: Regular Business, followed by a film show.